

Extremely Low-Noise Cryogenic Amplifiers for Radio Astronomy: Past, Present and Future Marian W. Pospieszalski

Acknowledgment: All the past and present members of the CDL Amplifier Group





Outline

- Short history: important points in the development of cryogenic amplifiers at NRAO
- Review of the state-of the-art of cryogenic amplifiers with InP HFETs and SiGe HBTs.
- Cryogenic transistors and amplifiers: understanding of noise performance
 - Noise models; accuracy of model predictions
 - Dependence on bias, optimal noise bias: "quality of pinch-off"
 - Dependence on ambient temperature
 - Broadband noise matching
- On the limits of achievable noise performance of microwave transistors

A Short History of FET's

- 1952 Junction FET (JFET), Shockley (BTL)
- 1966 Schottky Gate FET (MESFET), Mead (Caltech)
- 1967 MESFET on GaAs, Hooper and Lehrer (Fairchild)
- 1970 Prediction of carrier accumulation at heterointerface, Esaki and Tsu (IBM)
 - First cryogenic experiments, Loriou *et al* (France NTC), 120 K at 1 GHz
- 1976 Cryogenic experiments at X-band ,Liechti *et al*. (HP), 60 K at 12 GHz
- 1978 Mobility enhancement in GaAs/AlGaAs demonstrated,
 U.S. Patent for HFET (HEMT, TEGFET, MODFET, SDHT), Dingle *et al.* (BTL)



A Short History of FET's (2)

 1980-82 – Practical cryogenic amplifiers using MESFET's demonstrated Weinreb *et al.* (NRAO), 7 K at 1.5 GHz, 20 K at 4.5 GHz





A Short History of FET's (2)

- 1980 demonstration of HFET (GaAs/AlGaAs) devices, Mimura *et al.* (Fujitsu), Morkoc *et al.* (U. of Illinois, Rockwell)
- 1984 -87 GaAs/AlGaAs HFET at cryogenic temperatures, Pospieszalski, Weinreb (NRAO), 6 K at 8.4 GHz



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A Short History of FET's (3)

- 1985 Pseudomorphic HFET (InGaAs/AlGaAs) introduced, Zipperian *et al.* (U.of Illinois), Rosenberg *et al.* (IBM)
- 1987 Pulse-doped PHEMT introduced, Moll *et al*. (HP)
 - First demonstration of InGaAs/InAlAs/InP HFET's, Morkoc *et al.* (U. of Illinois)
- 1988 Noise model of FET suitable for cryogenic applications developed, Pospieszalski (NRAO)
 - Pseudmorphic HEMT cooled, Weinreb *et al*. (NRAO),
 25 K at 40 GHz (Linear Monolithics)
 - First .1 um gate length InP HFET demonstrated, Mishra *et al.* (HRL) soon joined by TRW, GE and Martin-Marietta (1989-91)



8-18 GHz Amplifier at 12.5 K (1988)



First amplifier designed using "Pospieszalski " noise model

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VLBA Q-band Amplifier (1991)



Linear Monolithics – ROHM Research PHEMT with .1 X 100 micron gate has been used These amplifiers are still in use in majority of VLBA Q-band receivers



InGaAs/InAIAs/InP HEMT







State-of-the-Art 2016





JVLA Amplifiers





MIC ("chip and wire") vs. MMIC

MMIC Disadvantages:

- Package typically over-moded (absorbers needed)
- Repair of failures much more difficult,
- If it does not work properly, no diagnosis can be easily performed
- Impossible to modify
- Limited number of chips on a single wafer
- Performance not always repeatable from wafer to wafer

MMIC Advantages:

- Assembly labor (saves anywhere from 1 to 4 days of technician time per amplifier)
- Much better control of design than in case of MIC
- MIC practically impossible to built with devices having gates < 100 nm





What we do not understand about cryogenic HEMTs:

- light illumination: effects on noise and gain, short and long term stability of both
- 1/f gain fluctuations
- what determines the "quality of pinch off"
- DC (or possibly RF) instability in InP HFET's of larger gate periphery (>200 microns)
- so called "current switch off" effect in InP devices which seems to be dependent on device layout
- an additional source of noise with 1/f like spectrum at certain bias usually at drain currents densities slightly higher than typically optimal for noise



Best Cryogenic SiGe Amplifiers





What We Do Understand



Common Noise Representations of 2-Ports





Common Noise Representations of 2-Ports (2)





Allowed Values of Noise Parameters

For All Linear Noisy Two-Ports: $|\rho| \le 1 \quad \iff \quad T_{\min} \le 4NT_{o}$ $T_{\min} = T_{0} \{2N + Re(\rho \sqrt{R_{n}g_{n}})\}.$

If therefore Re(ρ) ≥ 0 and correlation matrix is Hermitian and non-negative definite, than always

 $1 \leq \frac{4 \,\text{NT}_0}{\text{T}_{\text{min}}} \leq 2$ For all microwave transistors for useful frequency range : $\frac{4 \,\text{NT}_0}{\text{T}_{\text{min}}} \approx 2$

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Simplest Noise Equivalent Circuit of a FET



Noise Parameters of FET: Low Frequency Approximation





Noise Parameters of FET: High Frequency Approximation



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Simplest Noise Equivalent Circuit of Intrinsic HBT



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InP HFET vs. SiGe HBT

In P HFET

$$T_{min} \cong 2 \frac{f}{f_t} \sqrt{g_{ds} T_d r_t T_g}$$

$$f \ll \frac{f_t}{\sqrt{\beta}} \text{ and } \beta \gg 1$$

$$T_{min} = \sqrt{\frac{2 qI_b T_a r_b}{k}}$$

$$f \gg \frac{f_t}{\sqrt{\beta}} \text{ and } \beta \gg 1$$

$$T_{min} = \frac{f}{f_t} \sqrt{\frac{2 qI_c T_a r_b}{k}}$$

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Accuracy of Noise Model





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ALMA Band #1 5 Stage Amplifier: Modeled and Measured Results at 20 K







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Bias Dependence



Example of T_g and T_d Dependence on I_d



Optimal Noise Bias of GaAs FET at 297 K (1991)





Optimal Noise Bias of InP HFET at 18 K (1994)



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Example of Equivalent Circuit and $g_m(I_{ds})$ Characteristics of Cryo3 InP HFET



T_{min} Dependence on Bias (Summary)

$$T_{\min} \cong 2\frac{f}{f_t}\sqrt{g_{ds} T_d r_t T_g} \cong \frac{f}{f_{max}}\sqrt{T_d T_g} \qquad f_t \cong \frac{g_m}{2\pi(C_{gs} + C_{gd})}$$

Noise optimal bias is minimizing the value of: $f(V_{ds}, I_{ds}) \approx \frac{1}{2}$

The importance of gate recess technology (Joel Schleeh, Chalmers/LNF)



WAMICON 2015, Tutorial on Noise in National Radio Science Meeting, Jan. 4-6, 2017



ds

 $g_{\rm m}$

Cryogenic T_{min} at 8.4 GHz and dc pinch–off charactersitcs of GE HFET's (1987)





Ambient Temperature Dependence



S-Parameters of K-Band Amplifier Versus Ambient Temperature



RWW 2017



Noise Temperature of K-Band Amplifier Versus Ambient Temperature







Broad Band Noise Matching



Any FET or Bipolar Transistor is practically a THREE Noise Parameter Device



For a given frequency range chose Γ_{opt} close to SC center.

In practice, for a given frequency range average T_n is:

$$\frac{1}{f_{\max} - f_{\min}} \int_{f_{\min}}^{f_{\max}} T_n df \approx M_{\min}(f_{\max})$$

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Measured and Modeled T_{min} of a FET







Measured and Modeled $4NT_0/T_{min}$ of a FET











Broad Band Noise Matching Illustration (1)





Broad Band Noise Matching Illustration (11)



Low Noise Factory Catalog Data: http://www.lownoisefactory.com/ ALMA Memo #601, 2016



Device Scaling: Gate Length

$$T_{\min} \cong 2 \frac{f}{f_t} \sqrt{g_{ds} T_d r_t T_g} \cong \frac{f}{f_{max}} \sqrt{T_d T_a}$$

$$L_g \qquad f_{max} \swarrow T_{min} \qquad if T_d \approx const.$$

But within a measurement error no device demonstrated T_{min} lower than that predicted 25 years ago. The best cryogenic wafers: Chalmers (130 nm), NGSTCryo3 (80-100 nm), NGST (35 nm) exhibit progressively better f_{max} and M_{min} but about the same minimum T_{min} because T_d increases for deep submicron gate lengths.

General and (Very Simple) Picture of Noise in FETs:



 Γ^2 shot noise suppression factor

 Γ^2 should:

- be approximately independent of FET bias and its physical temperature.
- for long gates Γ² should assume a constant value while for short gates it should increase as in the limit for L_g→0, a pure shot noise should be observed and Γ²→1.
- As the average energies of hot electrons in Si, GaAs and InGaAs which form channels of all modern FETs are not that different for electric fields larger than 10⁴ V/cm (1 eV), Γ² should be only weakly dependent on a particular semiconductor structure





















Concluding Remarks

- Only three wafer runs of InP discrete devices (NRAO/HRL, WMAP/HRL, NGST/JPL cryo3) have been used in construction of great majority of radio astronomy instruments: VLA/EVLA,VLBA, GBT, ALMA band6, CBI, SZ-Array, WMAP, Planck LFI (K_a and Q), VSA, AMI, MPI, JPL/DSN and others.
- MMICs with very competitive noise are now becoming available.
- There has been no significant progress in the low noise performance of cryogenic HFET's in the past 15 years, as we are approaching the limits determined by the principle of operation
- Amplifier noise temperature is no longer the dominant component of the system noise for radio astronomy instruments with cryogenic receivers
- SiGe HBT offer competitive noise temperatures only at low frequencies but offer significantly better 1/f gain variations.



State-of-the-Art 20....?







www.nrao.edu science.nrao.edu public.nrao.edu

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SiGe HBT



Beta increases from 300 to 8000!

Courtesy of J.Bardin University of Massachusetts

Gm increases by factor of 3!



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Extracted f_t of SiGe HBT

Courtesy of J.Bardin University of Massachusetts

Typical peak f_t increase of 50% observed with cooling to 18 K!



18K 50K 77K 200K 300K











"Hot electron" noise



After Fischetti, IEEE Trans. ED vol. 38, p.634, March 1991

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