



## Implementation Status of the Ultra-Wideband Receiver Package For The North American Array

Dr. Jose Velazco Applied Electromagnetics Group Jet Propulsion Laboratory

2017 USNC–URSI National Radio Science Meeting 4-7 January 2017 Boulder, Colorado, USA







## Acknowledgements

• Collaborators:

Melissa Soriano, Dan Hoppe, Damon Russell, Ezra Long, Jim Bowen, Lorene Samoska, Andrew Janzen, Larry D'Addario

 This work is being carried out with funding from JPL's Research and Technology Development Program



Outline



- 1. Overview
- 2. System Requirements
- 3. Development Status
  - a. Feed horn
  - b. Low-noise amplifier
  - c. Cryogenic package
- 4. Summary









# 2. System Requirements

#### System Requirements

- $\odot$  Continuous coverage for 8-48 GHz
- $\circ$  T<sub>rx</sub> < 30 K,
  - T<sub>svs</sub> of 34 K @ 10 GHz and 45 K @ 30 GHz required
- $\circ$  Dual polarization

#### **Design considerations**

- Easy to manufacture
- $\odot$  Easy to service
- Compact cryogenic package
- Low Cost

#### Assumptions

- Baseline antenna is offset Gregorian (e.g. MeerKAT antenna) scaled to 18 m diameter with f/D = 0.55
- $\circ$  Optimize for 10 GHz



2. System Requirements



Implementation Status of the Ultra-Wideband Receiver Package For The North American Array



Ultra-wideband receiver package is composed of:

- O Quad Ridge Feed Horn (QRFH)
- Low-noise amplifier
  - Down-conversion stage

→ Task plan is to construct and test this receiver package

Note: Notional layout shown, final layout will be more compact





## 3.a. Status - Feed Horn

- Offset Gregorian paraboloid antenna assumed
- 8-48 GHz operation
- 0.55 focal ratio (similar to MeerKAT)
- Quad Ridge Feed Horn (QRFH)
  - Builds upon 6:1 design developed initially by A. Akgiray[1]
- Dielectric rod inserted to improve efficiency
  - Similar approach used by ATNF for Parkes Telescope ultra-wide band feed
  - Center radius of rod roughly determines coupling from horn to the rod vs. frequency, length determines pattern beam width (gain)
  - [1] Akgiray, A. H. (2013) Ph.D. Dissertation, California Institute of Technology



#### Jet Propulsion Laboratory California Institute of Technology

## 3.a. Status - Feed Horn

- Quad Ridge Feed Horn (QRFH) developed in collaboration with Sandy Weinreb's group
- Optimized for f/D = 0.55.



Implementation Status of the Ultra-Wideband Receiver Package For The North American Array



 Simple tapered Teflon rod is being modeled to demonstrate improvement in QRFH efficiency



#### Metal Only Horn with Dielectric Rod



## 3.a. Status - Feed Horn

Best Rod Parameters:  $R_2=2.3 \text{ mm}$   $L_1=L_3=5 \text{ mm}$  $L_2=20.4 \text{ mm}$ 

- Center radius of the rod roughly determines the coupling from the horn to the rod vs. frequency
- Rod length determines the pattern beam width (Gain)









#### 3.a. Status - Feed Horn

Implementation Status of the Ultra-Wideband Receiver Package For The North American Array





#### **Quad Ridge Feed Horn**

![](_page_10_Picture_0.jpeg)

![](_page_10_Picture_3.jpeg)

- 3.b. Status LNA Design
  - Designing custom MMICs, getting them fabricated at foundries, installing in connectorized packages, and testing the packaged amplifiers
  - We are pursuing two LNA design approaches:
    - 70-nm GaAs m-HEMT from OMMIC
    - 35-nm InP HEMT from NGC

![](_page_11_Picture_0.jpeg)

![](_page_11_Picture_3.jpeg)

## 3.b. Status - LNA Design

![](_page_11_Picture_5.jpeg)

Preliminary experimental results obtained with **70 nm GaAs m-HEMT** from OMMIC in the **1-18 GHz** 

- 35 dB gain @ 8.4 GHz
- 5 K noise @ 8.4 GHz

Experimental LNA test results at 1-18 GHz range

![](_page_12_Picture_0.jpeg)

## 3.b. Status - LNA Design

LNA design for 8-48 GHz operation,

 70 nm GaAs m-HEMT from OMMIC

40 35 30 Gain 25 20 15 ¥.10 Noise 5 Input Match 0 -5 -10 -15 Output Match 15 50 10 20 25 30 35 40 50 GHz

Implementation Status of the Ultra-Wideband Receiver Package For The North American Array

![](_page_12_Picture_6.jpeg)

#### Wideband LNA Requirements

Parameter		Requirement
Noise Temperature	8-40 GHz	≤ 12 K
	40-48 GHz	≤20 K
Gain		≥ 30 dB
Gain Flatness		≤ 6 dB
Input Match	8-15 GHz	≤ -5 dB
	15-48 GHz	≤ -10 dB
Output Match		≤ -10 dB

![](_page_12_Picture_9.jpeg)

#### Simulation of OMMIC Wideband LNA

![](_page_13_Picture_0.jpeg)

![](_page_13_Picture_3.jpeg)

## 3.b. Status - LNA Design

 Earlier InP HEMT (NGC) testing at higher frequencies (40--300 GHz) showed outstanding performance, suggesting similar performance is possible below 40 GHz

![](_page_13_Figure_6.jpeg)

Measured results for 35 nm InP HEMT MMICs (NGC) over 40-300 GHz, with record noise temperature throughout

![](_page_14_Picture_0.jpeg)

![](_page_14_Picture_2.jpeg)

#### 3.b. Status - LNA Design

 Design of 8-48 GHz wideband LNA housing nearly completed

![](_page_14_Figure_5.jpeg)

![](_page_15_Picture_0.jpeg)

![](_page_15_Picture_2.jpeg)

#### **3.c.** Status – Cryo-package

- Design has been implemented in SolidWorks
- Currently performing thermal design analysis for ~5 K operation
- Quantum GA1 cryogenic system purchased for evaluation and comparison with Sumitomo/CTI

![](_page_15_Picture_7.jpeg)

![](_page_16_Picture_0.jpeg)

![](_page_16_Picture_2.jpeg)

#### **3.c. Status – Cryo-package**

![](_page_16_Picture_4.jpeg)

![](_page_17_Picture_0.jpeg)

![](_page_17_Picture_2.jpeg)

### 3.c. Status – Cryo-package

![](_page_17_Figure_4.jpeg)

![](_page_17_Picture_5.jpeg)

![](_page_18_Picture_0.jpeg)

![](_page_18_Picture_3.jpeg)

#### **3.c.** Status – Cryo-package

#### 3-D model (1/2 scale)

![](_page_18_Picture_6.jpeg)

![](_page_18_Picture_7.jpeg)

![](_page_19_Picture_0.jpeg)

4. Summary

![](_page_19_Picture_2.jpeg)

- Test and evaluate performance of first OMMIC wafer run
- Second wafer run with revised LNA design scheduled for April
- Optimize design for QRFH with dielectric rod and fabricate feed
- Evaluate performance of Quantum GA1 cryogenic system, finalize cryogenic system design.
- Work on downconverter design
- Prototype receiver expected to be complete in September 2017. It will then be tested on an existing antennas comparable to the ngVLA antenna.