Full-Mueller Mosaic Imaging with ALMA Jan. 5th 2018, Boulder, CO



Impact of in-beam (PB) effects

S. Bhatnagar



P. Jaganathan, U. Rau, ALMA Polarization Team (P. Cortez, S. Kameno-san, C. Hull, T. Hunter, C. Brogan) + B. Kirk

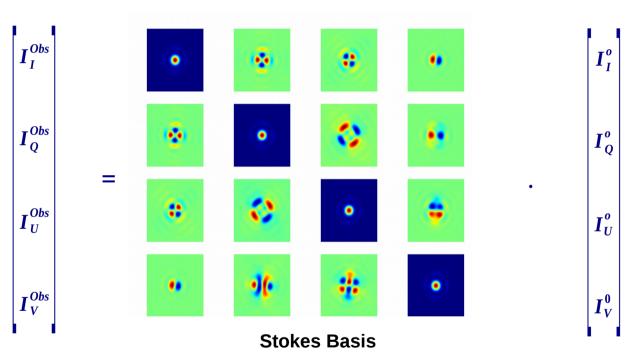
Plan for the talk

- Characterization of the problem
 - Theory of Full-Mueller imaging
 - Effect of PB: Apparent sky-polarization varies with time, direction and frequency
 - Effect of PB: Limits full-Stokes mosaic imaging performance
- Results from investigations with the EVLA
 - Effect of in-beam leakage on WF Stokes-Q, -U and -V imaging
 - Effect of Parallactic Angle (PA) coverage
 - Full-Mueller (WB) A-Projection
- ALMA Study Project: Full-Mueller Mosaic Imaging
 - Goals
 - Plan for the work and current status



Full-Mueller (-Polarization) WF Imaging

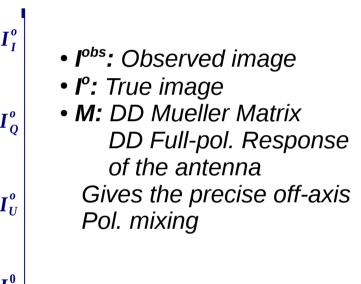
• Direction Dependent (DD) imaging equation



 $\boldsymbol{I}^{Obs} = [\boldsymbol{M}] \cdot [\boldsymbol{I}^{o}]$

- Diagonal: "pure" poln. products
- Off-diagonal: poln. leakage

 $M_{ii} = E_i \otimes E_i^*$





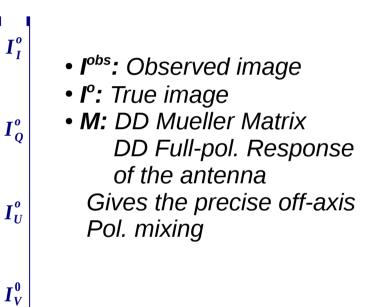
S. Bhatnagar: ALMA2030, URSI-NRSM, Jan. 4-6, 2018, Boulder

Full-Mueller (-Polarization) WF Imaging

• Direction Dependent (DD) imaging equation

 $I^{Obs} = [M] \cdot [I^o]$

- Diagonal: "pure" poln. products
- Off-diagonal: Include poln. leakage



 $\boldsymbol{M}_{ij} = \boldsymbol{E}_i \otimes \boldsymbol{E}_j^*$

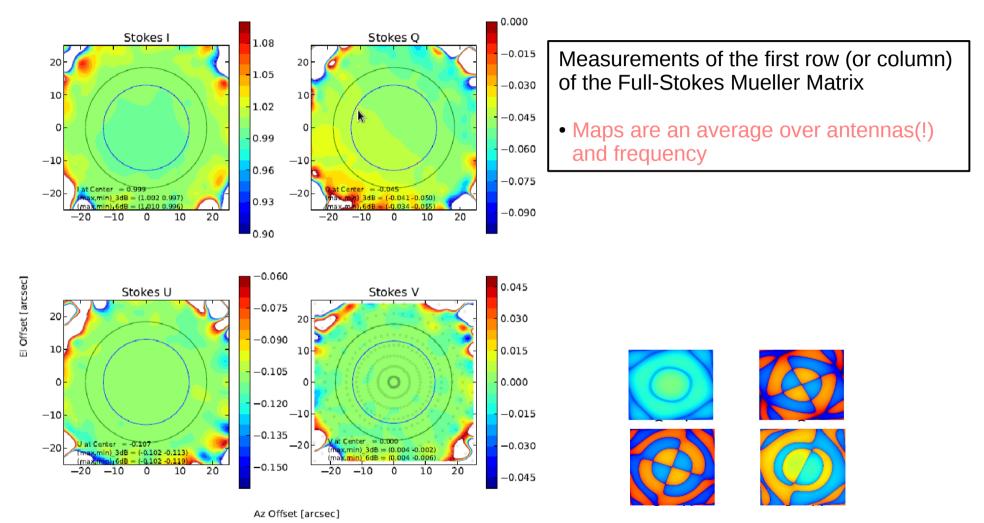
• Affects fidelity at the 10³⁻⁴ level



• PB Stokes-Q, -U is few% of Stokes-I

[Jagannathan, Bhatnagar, Rau & Taylor, AJ, 2017]

Measured full-Stokes PB maps



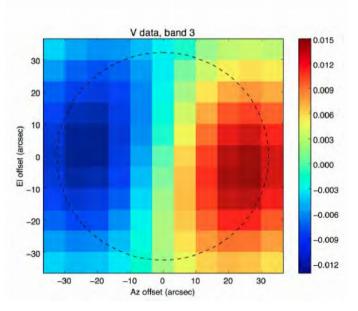
Expected from antenna opticsAlso measured for EVLA antennas

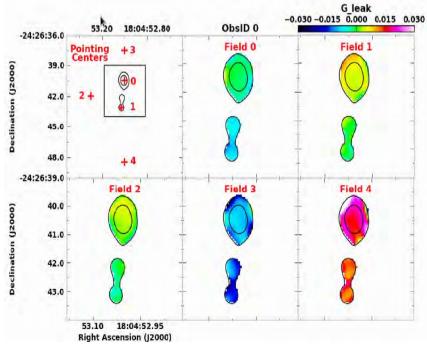
[S. Kameno & ALMA Pol. Team; 2015]



On-sky tests: Measured Stokes-V

- Measurement of Stokes-V for a CN(1-0) maser line emission
 - The measurements of the Stokes-V qualitatively match the <u>average</u> Stokes-V beam (Hull, 2015)
 - Does not match quantitatively (more leakage than shown by PB alone)
 - Does not reproduce the residual Stokes-V for <u>different times or with</u> <u>small offsets</u> (position in the beam) after taking into account the PB







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[Brogna, Hunter & Moellenbrock; 2015]

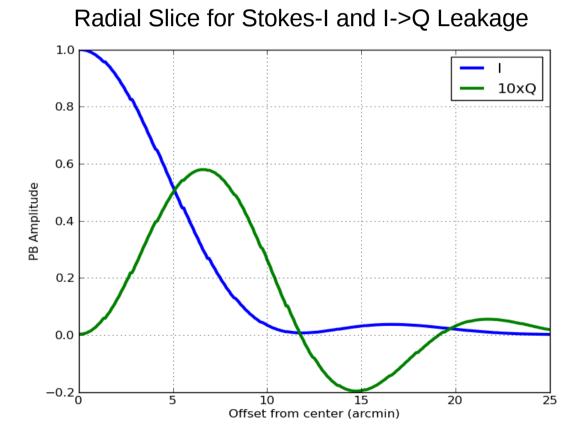
Full-pol. Imaging: Mosaic Sensitivity Pattern

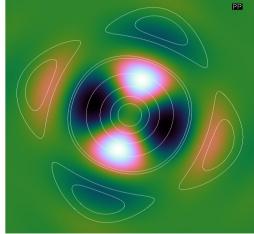
In-beam Stokes-Q pattern for a 11x11 point mosaic

- Pol. Leakage requriement for ALMA too loose to allow precision polarimetry ("...<3% within inner 1/3 PB").
- In-beam DD leakage pattern spreads across the mosaicked region
- Effects ignored here:
 - Heterogeneous case
 - Rotation due to PA change
- The resulting pattern is combination of overlapping Clover-leaf pattern of each pointing



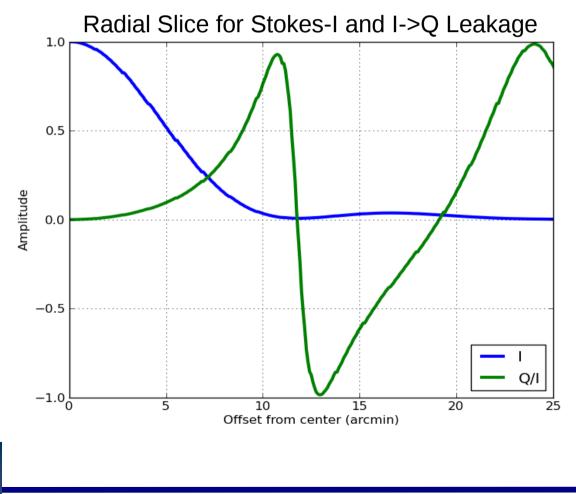
- Leakage (Off-diagonal elements of the Mueller matrix)
 - Vary with direction (position in the beam), Parallactic Angle (time) and frequency

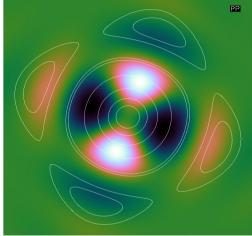






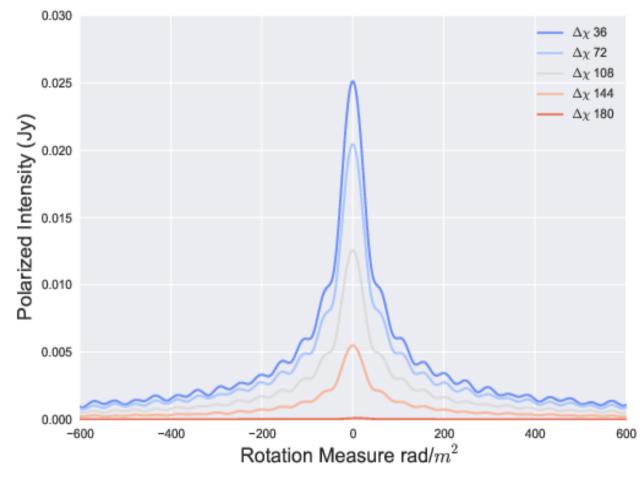
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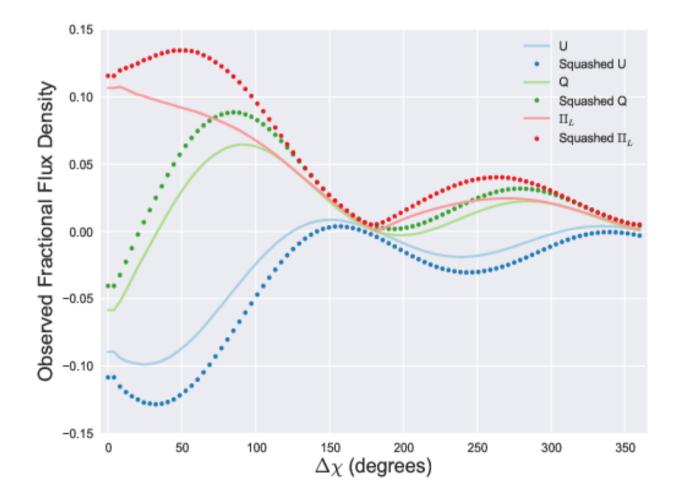
• Effect of PA averaging: Instrumental RM





[PhD Thesis, Jagannanthan, 2017] [Jagannathan, Bhatnagar, Rau & Taylo AJ, 2017]

• Effect of PA averaging: Residual instrumental leakage





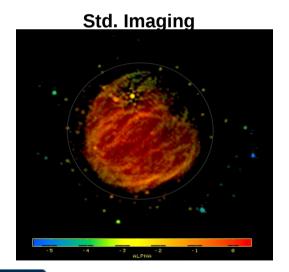
[PhD Thesis, Jagannanthan, 2017] [Jagannathan, Bhatnagar, Rau & Taylo AJ, 2017]

A-Projection algorithm

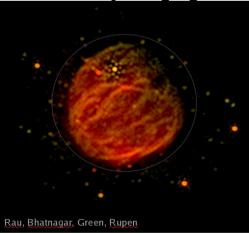
- Correct for the PB effects (DD effects in general) during imaging
- Essentially, apply a (pseudo) inverse of the Mueller matrix during imaging

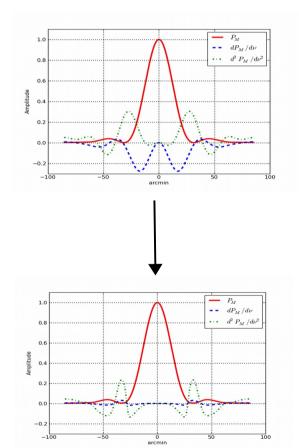
$$\boldsymbol{I}^{Obs} = \left[\boldsymbol{M}^{M}\right]^{-1} \cdot \left[\boldsymbol{M}\right] \cdot \left[\boldsymbol{I}^{o}\right]$$

Actual algorithms applies corrections in the visibility domain using antenna aperture illumination patterns (voltages).



A-Proj. Imaging





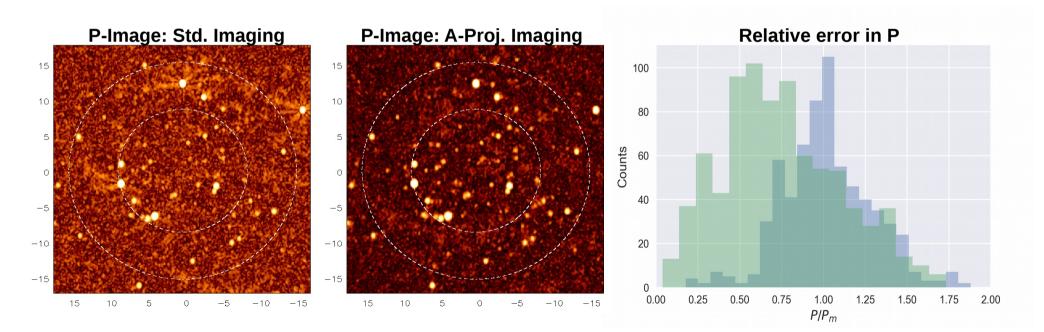


Needs good model for the PB (M^{M})

Full-Mueller A-Projection algorithm

- Correct for the PB effects (DD effects in general) during imaging
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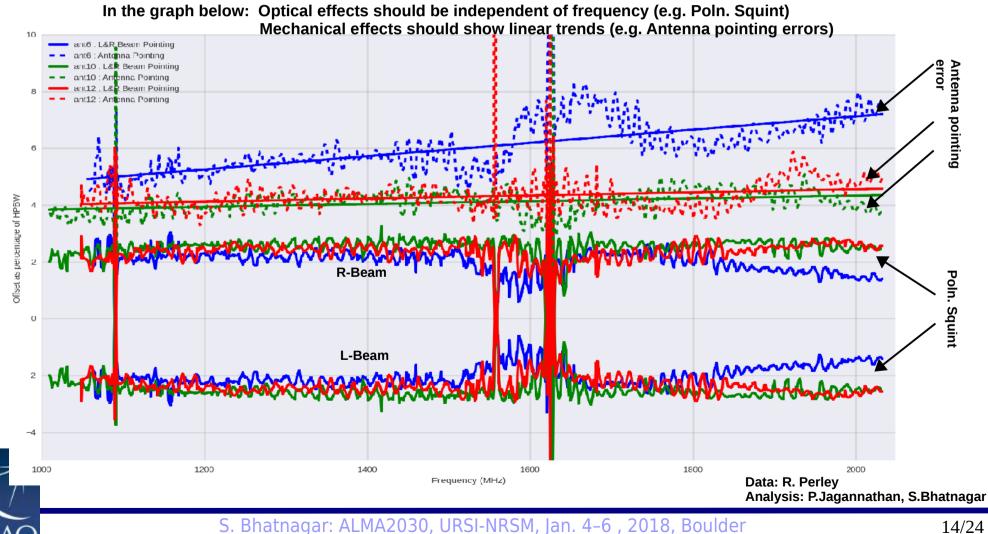


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[PhD Thesis, Jagannanthan, 2017]

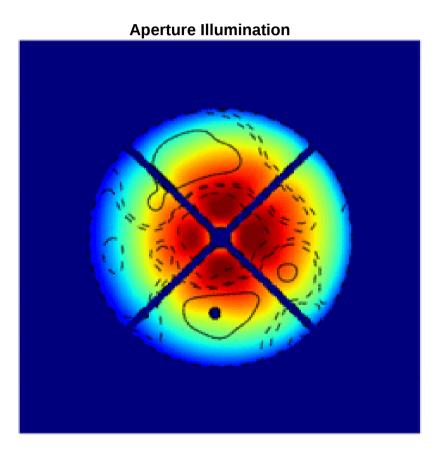
PB Effects: EVLA antenna holographic measurements

- Parametric model of antenna Aperture Illumination
 - Difference between Ant6 and Ant10 in a "homogeneous array" •



PB Effects: ALMA antenna holography

 Antenna-to-antenna variations in aperture illumination (older holography data)

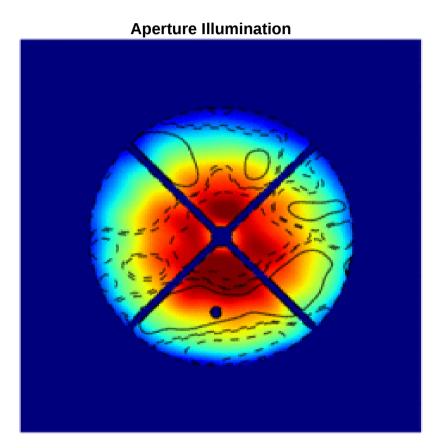


- Strongest effects
 - Antenna size, Pointing errors
 - Quadrapods
 - Antenna-to-antenna variations

[Kundert et al. IEEE Trans. A&P, V. 65, No.1, 2016]



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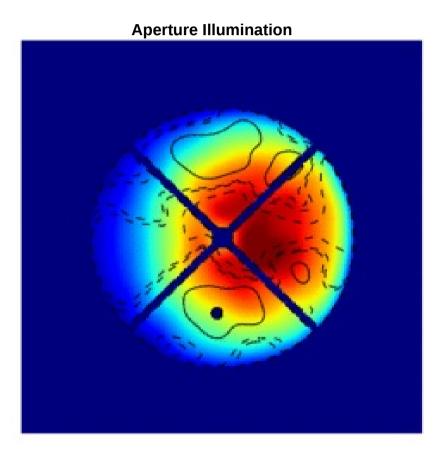


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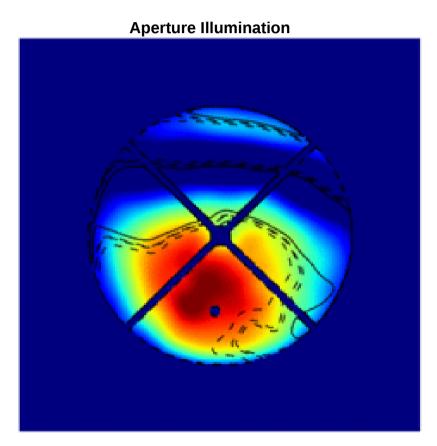


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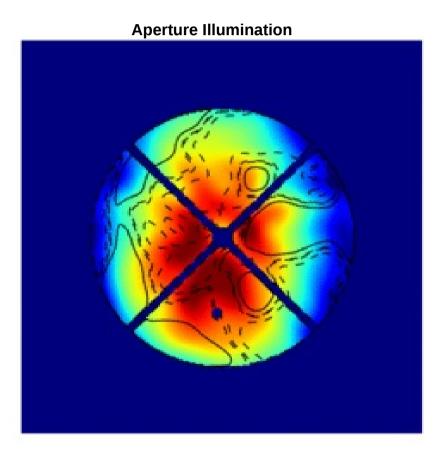


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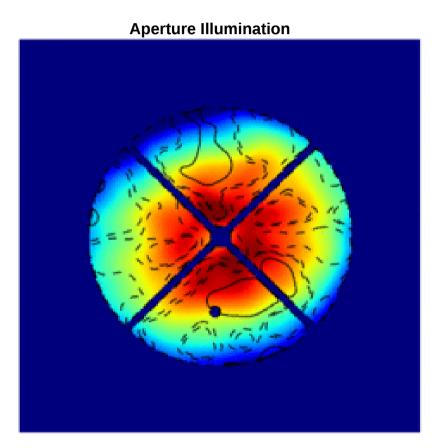


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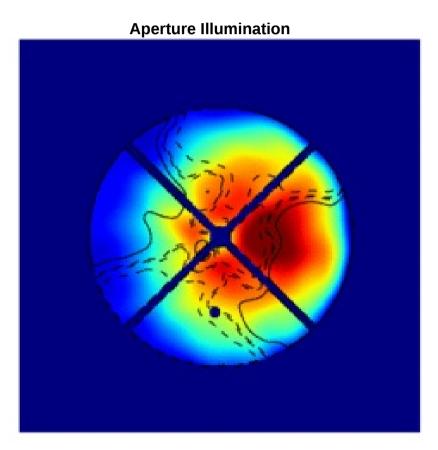


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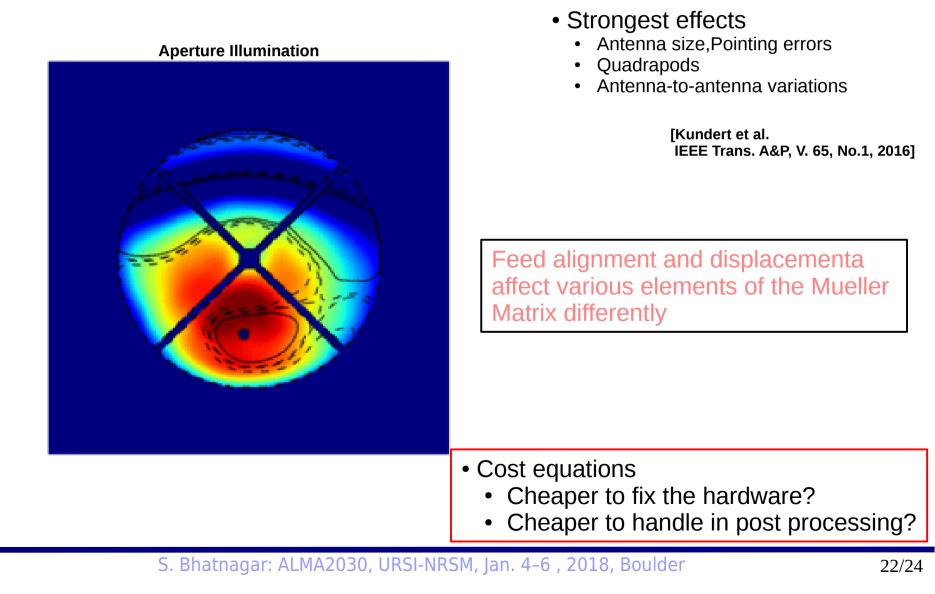


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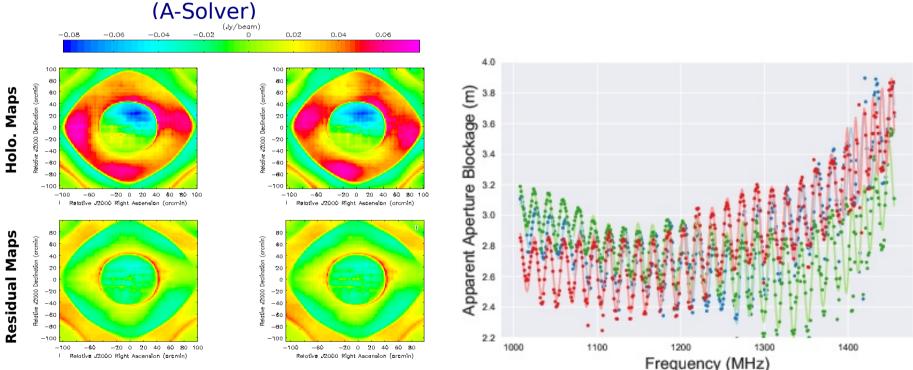




ALMA Study Project: Goals

Goals

1. Build a full-Mueller model for the antenna PB(s)



• Use the holography maps to solve for antenna structural parameters (A-Solver)

2. Incorporate the ray-trace models (heterogeneous array) in (WB) A-Projection for ALMA



Jagannathan, Bhatnagar & Brisken, 2016, AJ, 2017 [arXiv:1711.00875]

ALMA Study Project: Full-Mueller Mosaic Imaging with ALMA

- Plan for work
 - Understand the existing holography data
 - Zeeman effect (Stokes-V) measurements suggestive of residual PB effects
 - Possibly acquire new data using the artificial sources at the high site to get higher SNR PB maps.

Status

- Work started on Nov. 1, 2017
- Hired Brian Kirk
 - Former Senior-DA at CV, experienced in ALMA data processing and pipeline operations
- Re-processing data using standard (debugged) tools
 - DD Jones matrix (far field voltage patterns) per antenna as a function of frequency

